

Solar-B EIS Engineering Meeting a  
July 27 - 29, 1999  
H. Hara (NAOJ)

3/Solar-B

## Documents for Subsystem Design

Document Name	Date	Lang
Thermal Design Standard for Solar-B (DRAFT)	Feb 12,1999	J
Requirements for Structural Mathematical Models	Mar 8,1999	E
Mechanical Design Standard (DRAFT)	Mar 8, 1999	J+E
Environmental Conditions for Solar-B (DRAFT)	Mar 8, 1999	J+E
Contamination Control Program Plan for Solar-B (DRAFT)	Mar 8, 1999	E
Solar-B Electrical Design Standard (DRAFT)	Apr 24,1999	J
	(DRAFT) July 27,1999	E
Solar-B/Telescope Thermal Interface Condition (DRAFT)	June 30, 1999	J+E
Requirements for Interface Thermal Math Model of Solar-B Telescope (DRAFT)	June 30, 1999	J+E
Geometry & Command Design Standard (DRAFT)	not yet made	

/ report of satellite design meeting on June 30, 1999

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Solar-B

Requests from System side to EIS

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Position of Mounting I/F points. There is inconsistent with the I/F legs. Change the drawing to meet the I/F points before making structure math model.

Electrical relationship among ICU, MHC, and FPA. How these are connected electrical lines ?

Ranges of temperature is too narrow. Widen the ranges (relatively urgent).

Show disturbance torque of each moving component (urgent issue).

32 W solar power is input in EIS at entrance filter. Is this OK ? How is this energy to be treated ?

Number and location of survival heaters when EIS primary power is off. Temperatures for switch-on and switch-off controls should be reported. The survival heaters are controlled by HCE. At present three survival heaters are allocated to EIS by the request of Japanese side, though the usage is uncertain. This issue must be reported to the system side by July 15.

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## Solar-B system issues

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ometry & Command system: ASTRO-F system was adopted.

nary of electrical interface related to EIS was shown (see Fig. 1 ).

lite structure almost meet the size of fairing envelope.

option: XRT & EIS. In case of EIS, interface points should be shifted to -Z dir

is put at low part of +Y side panel of bus structure (Fig.2).

ht of EIS: EIS-STR 50.3 kg, EIS-ICU 6.0 kg, EIS-HAR 4.0 kg. Margin  
is managed by system side in a different level of system issue.

r: ICU:  $W_{\max} = 60 \text{ W}$ ,  $I_{\max} = 2.3 \text{ A}$

Heater power of 10 W is tentatively allocated to EIS because of no input.

nal analysis of optical bench + mounting legs + bus structure was almost finis  
ge of optical axes of SOT, XRT, and EIS is only 0.2-0.3 arcsec during the sun  
ronous orbit in the worst case. The infinite rigid body for telescopes is assum  
s analysis.

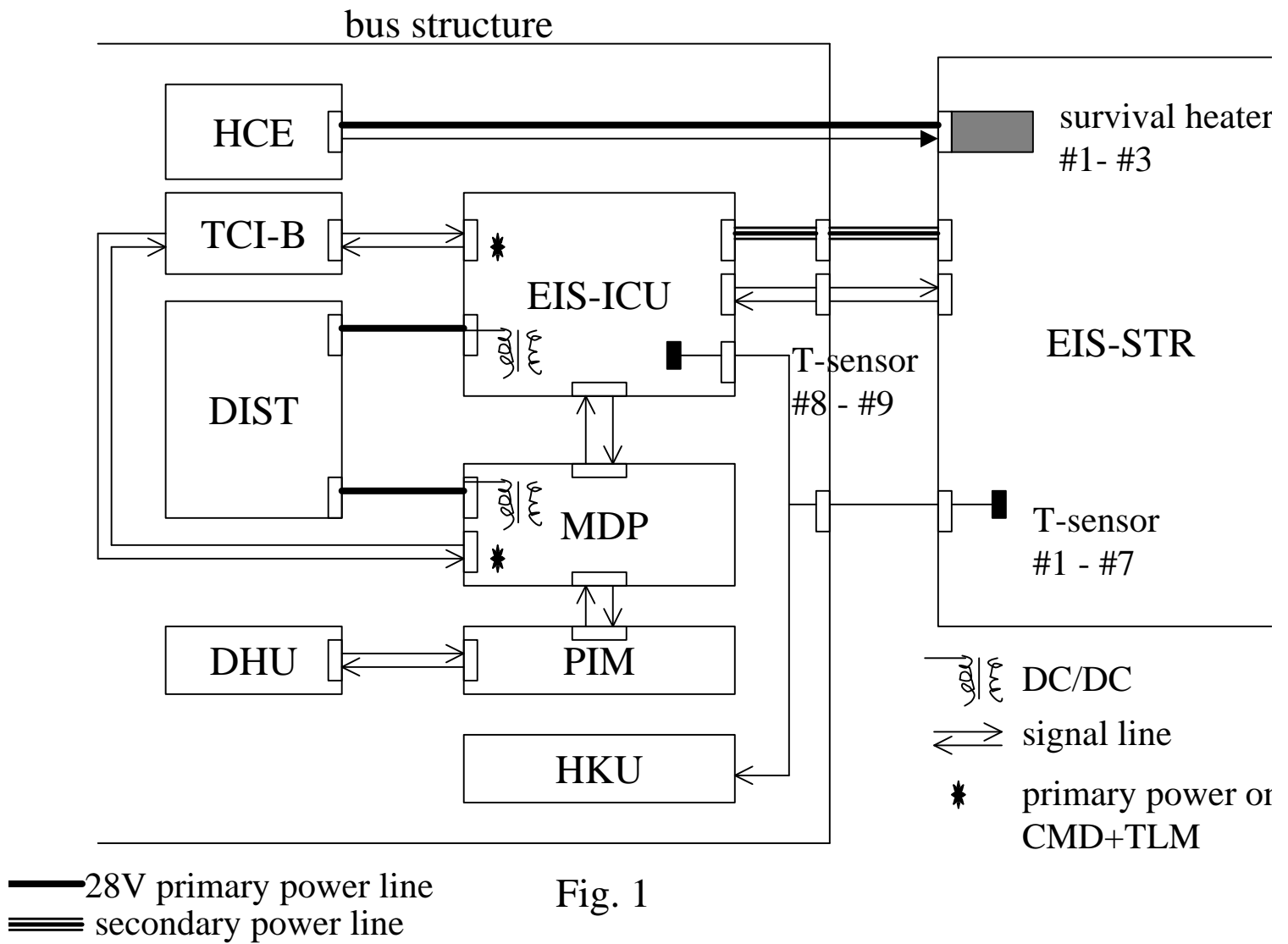
ss: If there is an electrical harness whose length should be less than 4 m by so reason, we must report it to the system side.

: savings: Even in the case of sun-synchronous orbit, satellite night comes in s period. System asked all telescope teams if the power saving can be during the night period or not for power saving.

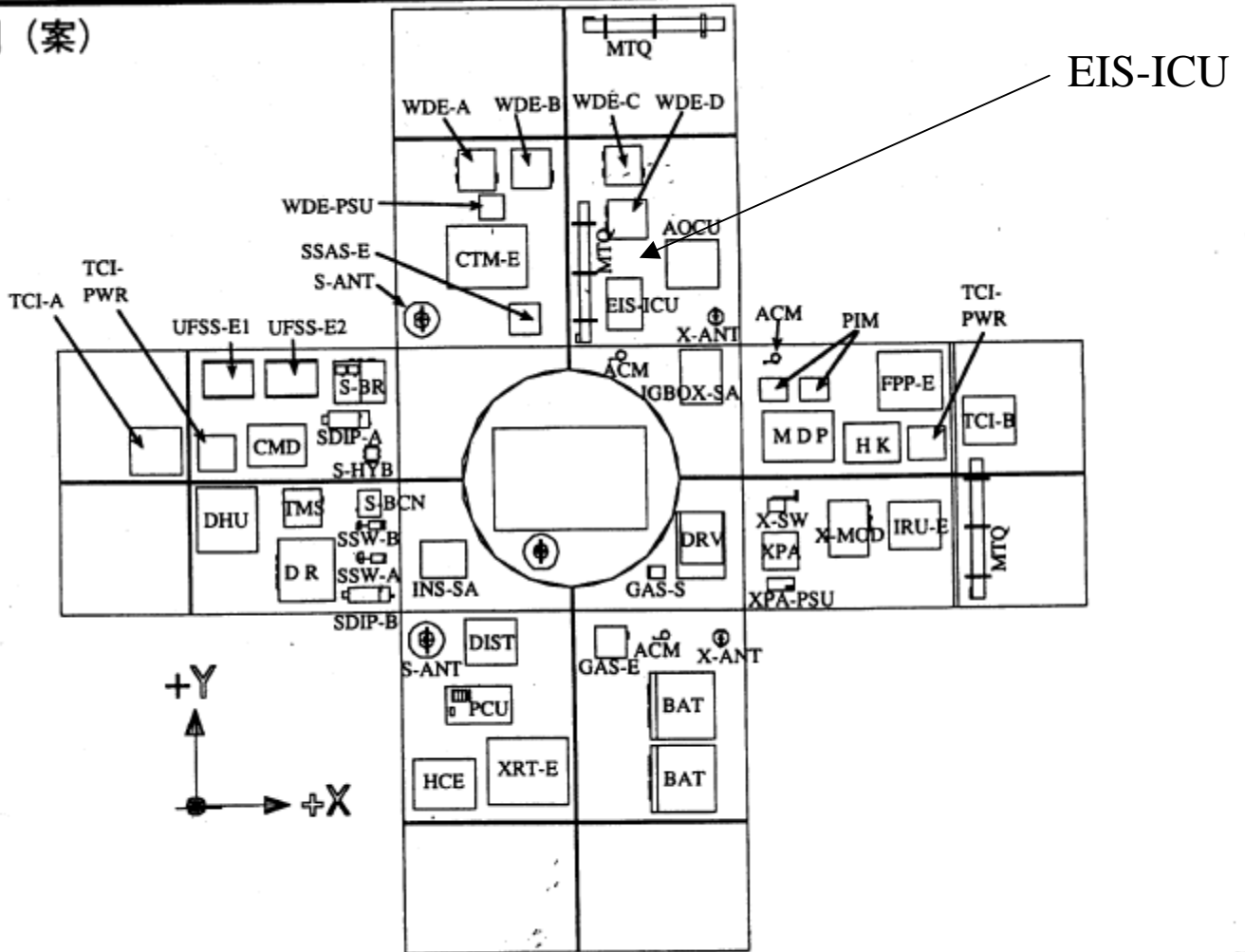
oltage for SOT, XRT, and EIS is  $28 \pm (A)$  V. (A) will be less than 2.

ure math model: The schedule of structure analysis was updated (see Fig. 3).

ial math model: The schedule of thermal analysis was updated (see Fig. 4).



器配置図 (案)



5.2

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Schedule of Structure Analysis

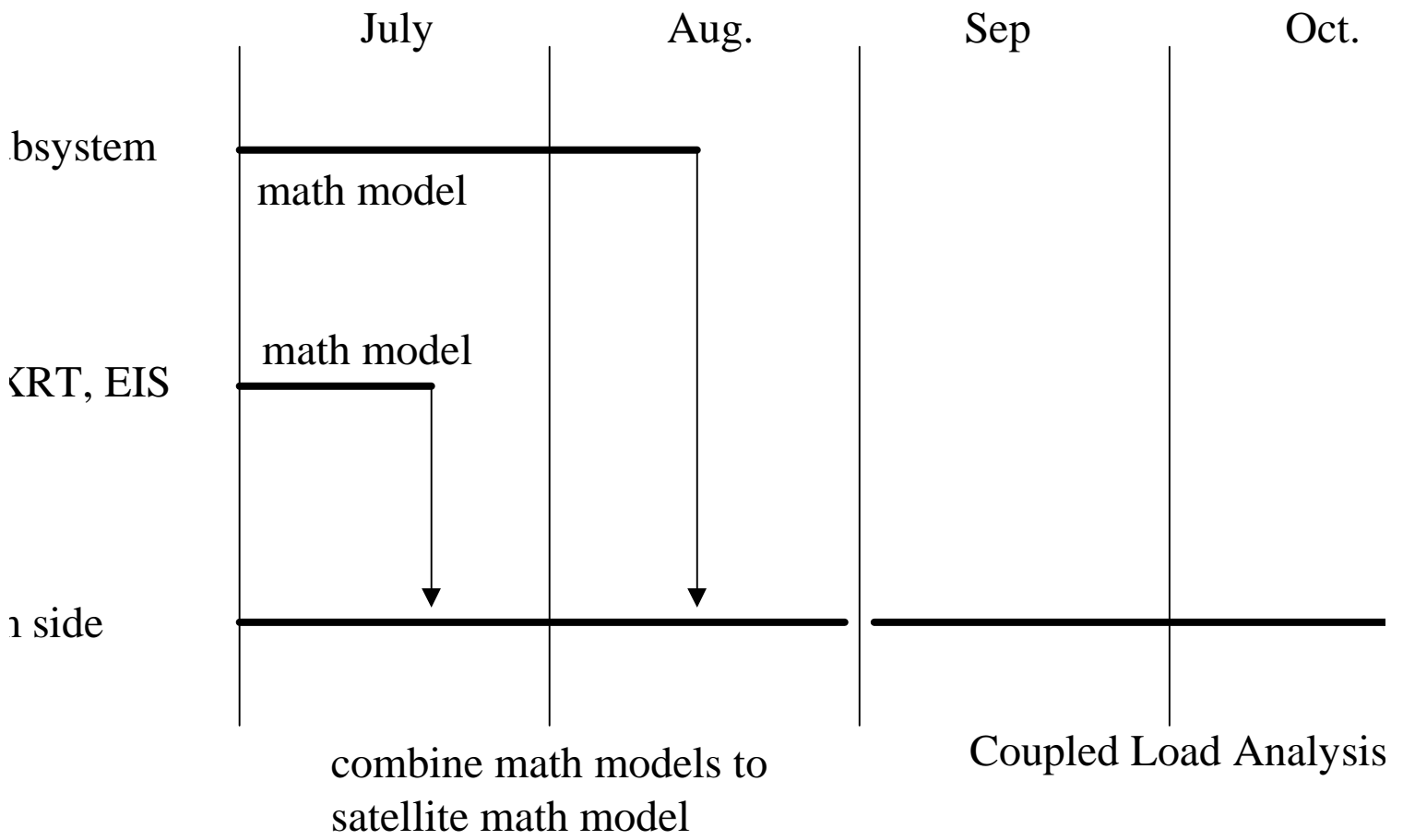
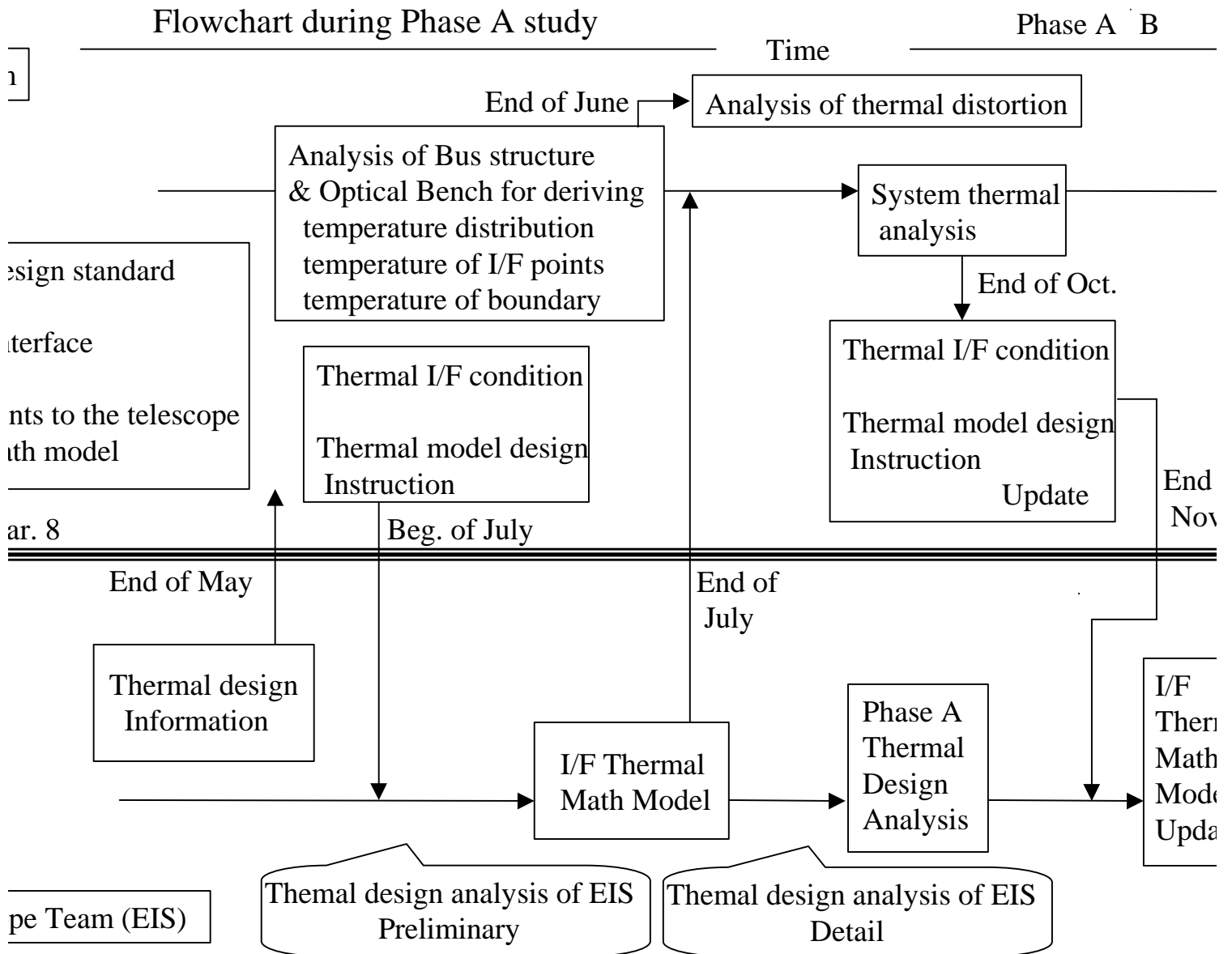


Fig. 3



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# Solar-B Schedule of Thermal Analysis



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## MDP development schedule

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Specifications of MDP shall be fixed by the following dates:

hardware related issues: August 31, 1999

software related issues: November 30, 1999

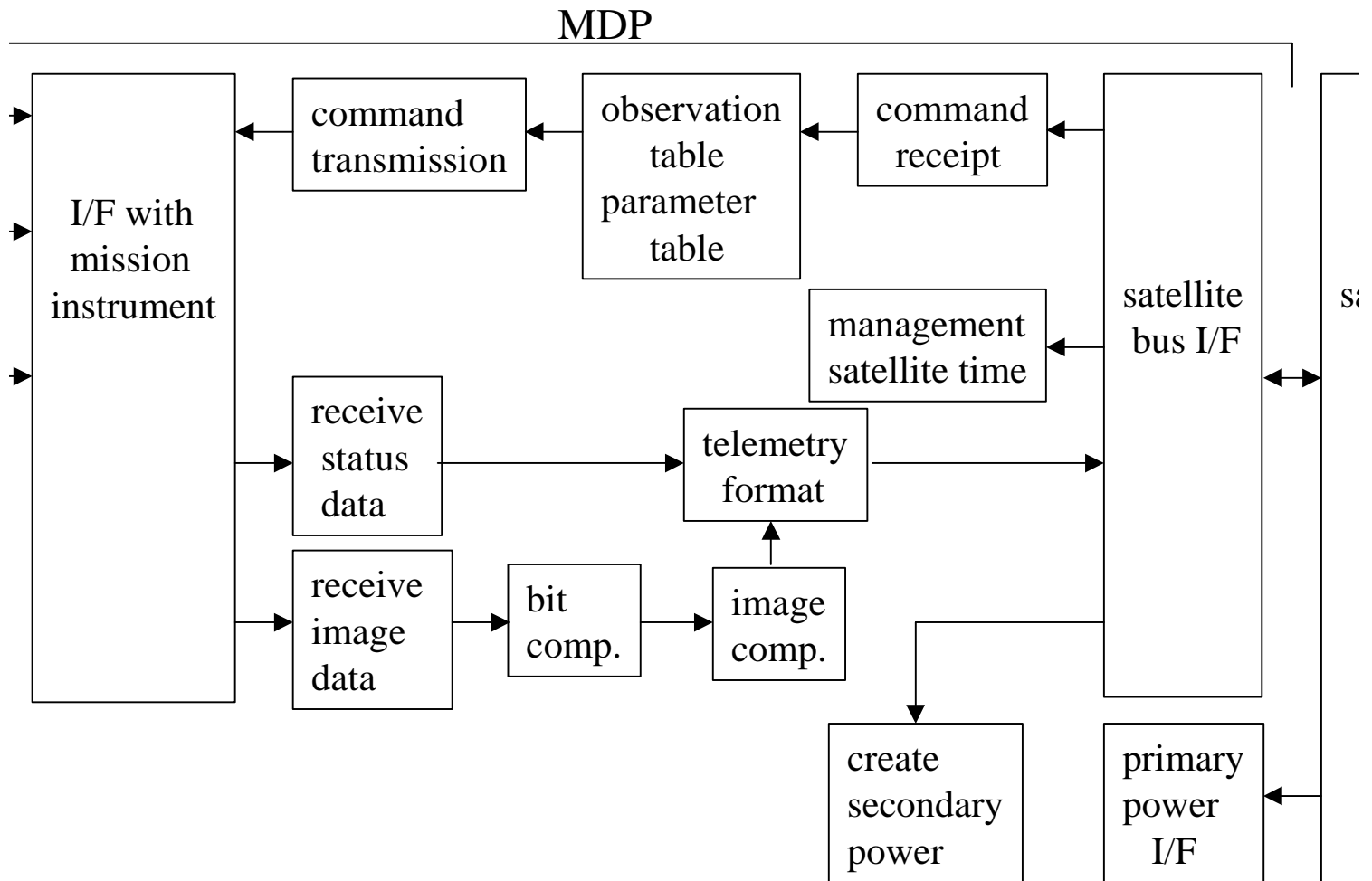
Purpose of PM tests (Feb 1 - Mar 31, 2001)

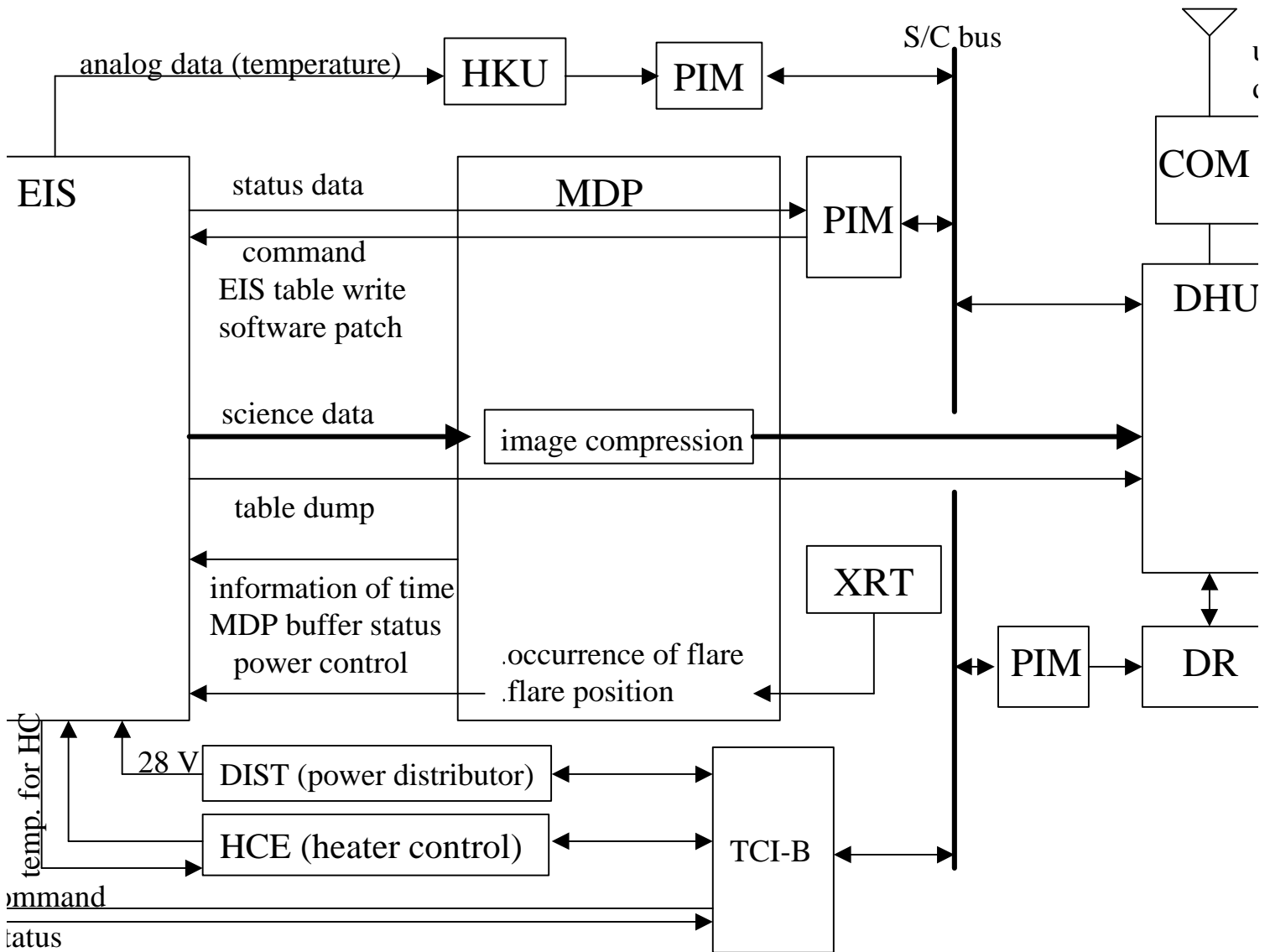
Interface with SOT, XRT, EIS, DHU, TCI-B, HCE, DIST

- establishment of hardware interface
- establishment of software protocol
- establishment of exposure sequence

Confirmation of main processing functions

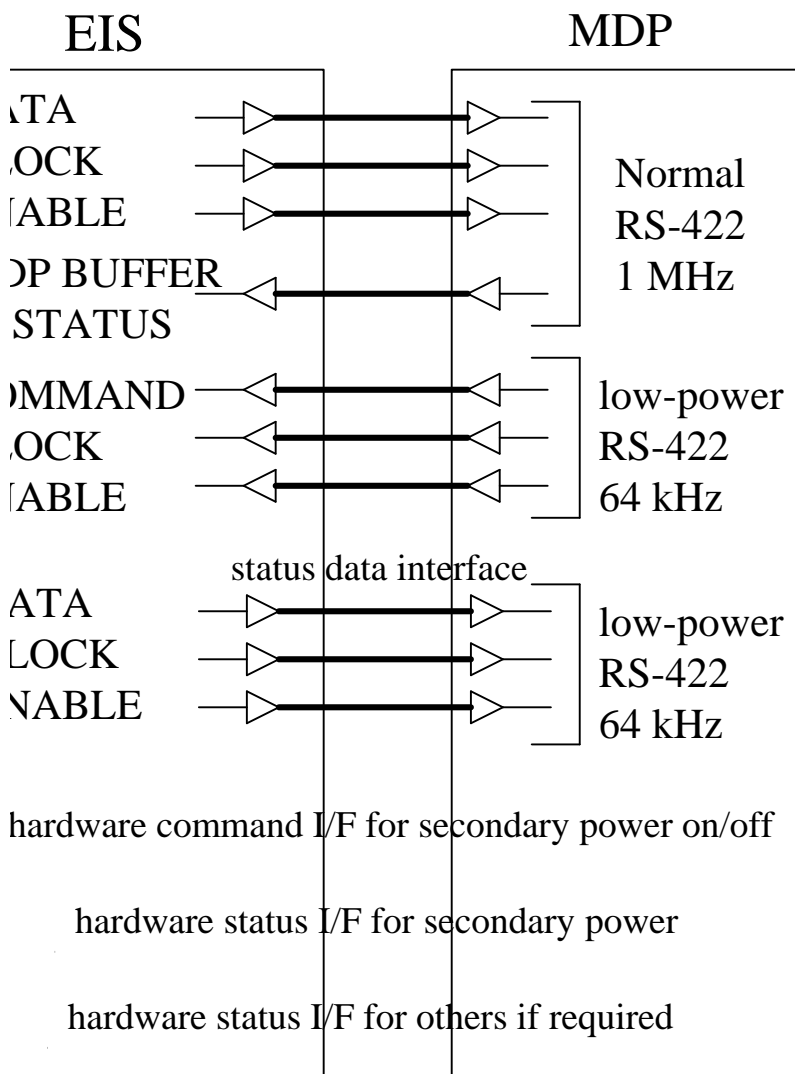
- confirmation of processing command/telemetry
- confirmation of processing science data including compressio



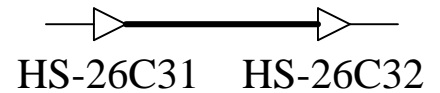


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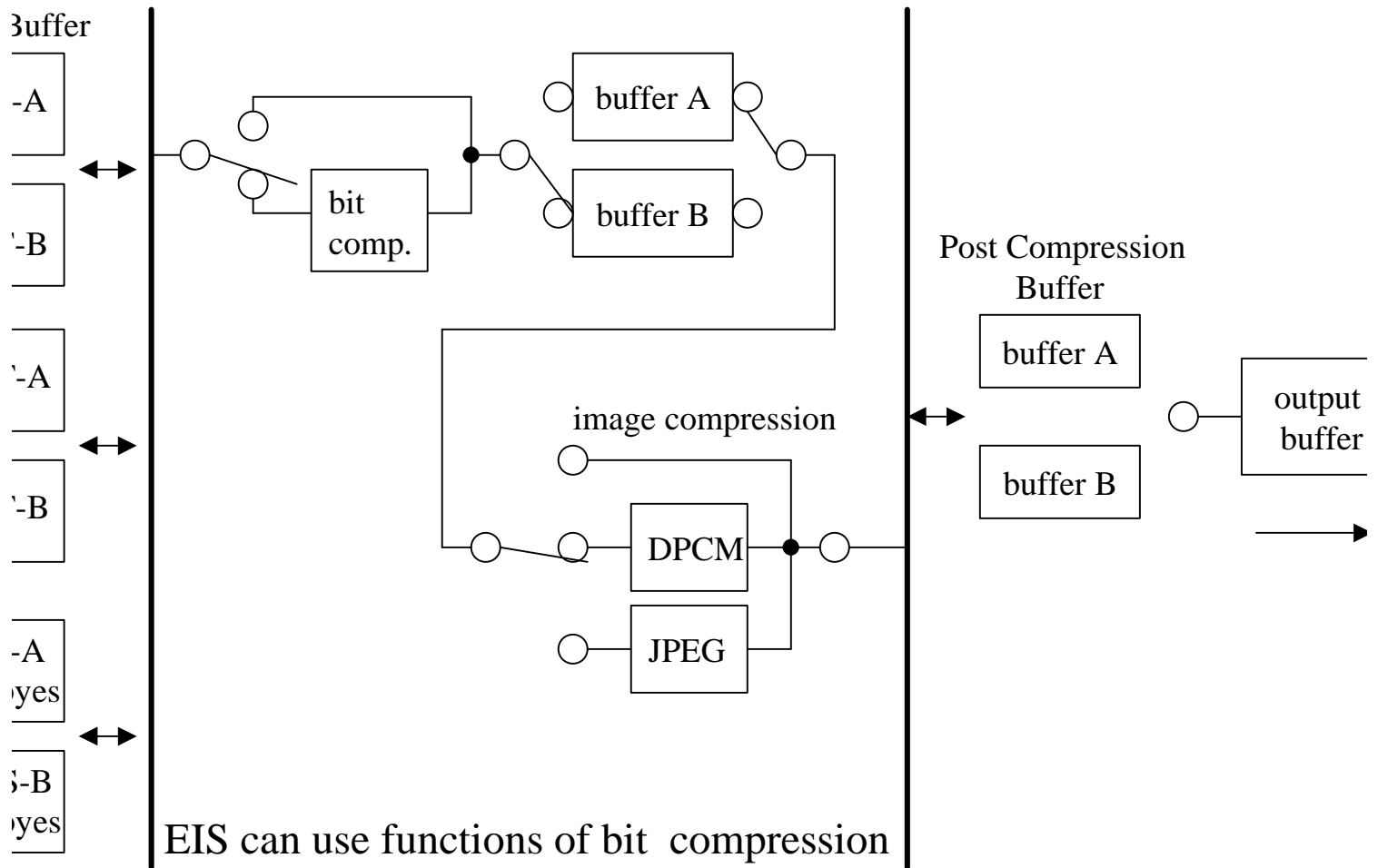
## Interface lines



All electrical interface lines between EIS and MDP are digital interfaces using differential RS-422 line interface



- ‘MDP Buffer Status line is used for:
- 1: for safe exchange of ‘input buffer’ in MDP.
  - 2: when post compression buffer become almost full.



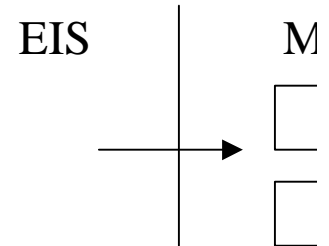
EIS can use functions of bit compression & image compression in MDP.

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## Data lines

line: normal differential RS-422 line, 1Mbps

- Science data (spectrum or image)
- Status data ( ~1 sec interval; Report of shutter & scan mirror motions will be required.)
- Dump data of EIS observation table
- Dump data of EIS compression table



MDP knows kind of data by looking at the header part after the data are transferred  
**2 bits data flow in this data line.**

maximum transfer speed: 1 Mbps (TBD)

Input Buffer in MDP: 1. double buffer system 2. buffer size: 8 Mbytes/buffer

concept of Data Format: fixed-length header + variable-length data

maximum image size in a single packet = 256 kpixel

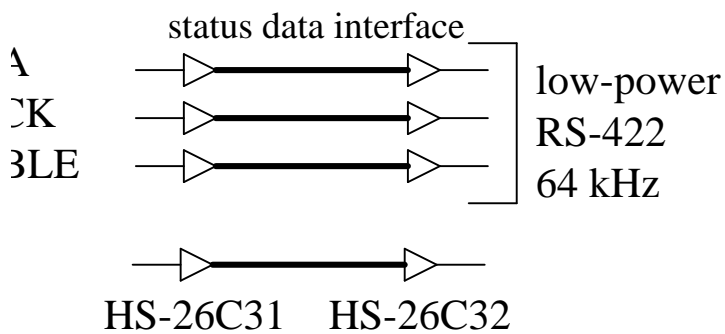
science data: header + 12 bits CCD image data

The number of pixels in the image data is a multiple of 8.

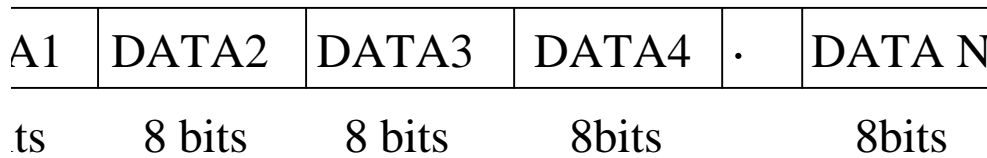
Is this OK ?

other data : header + 12 bit data (Upper 4 bits are all 0. Is this OK ?)

can send data to MDP at any time when a status of MDP buffer status line is R



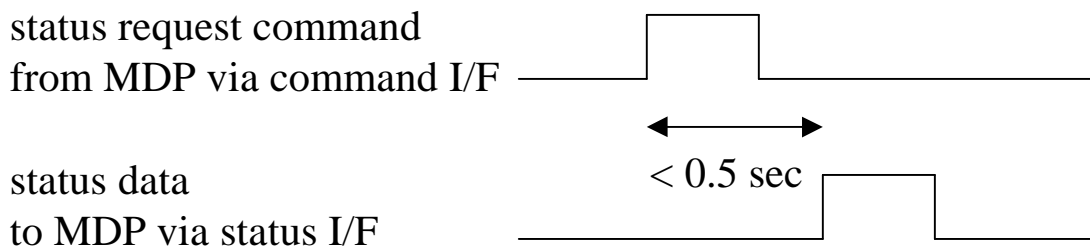
format



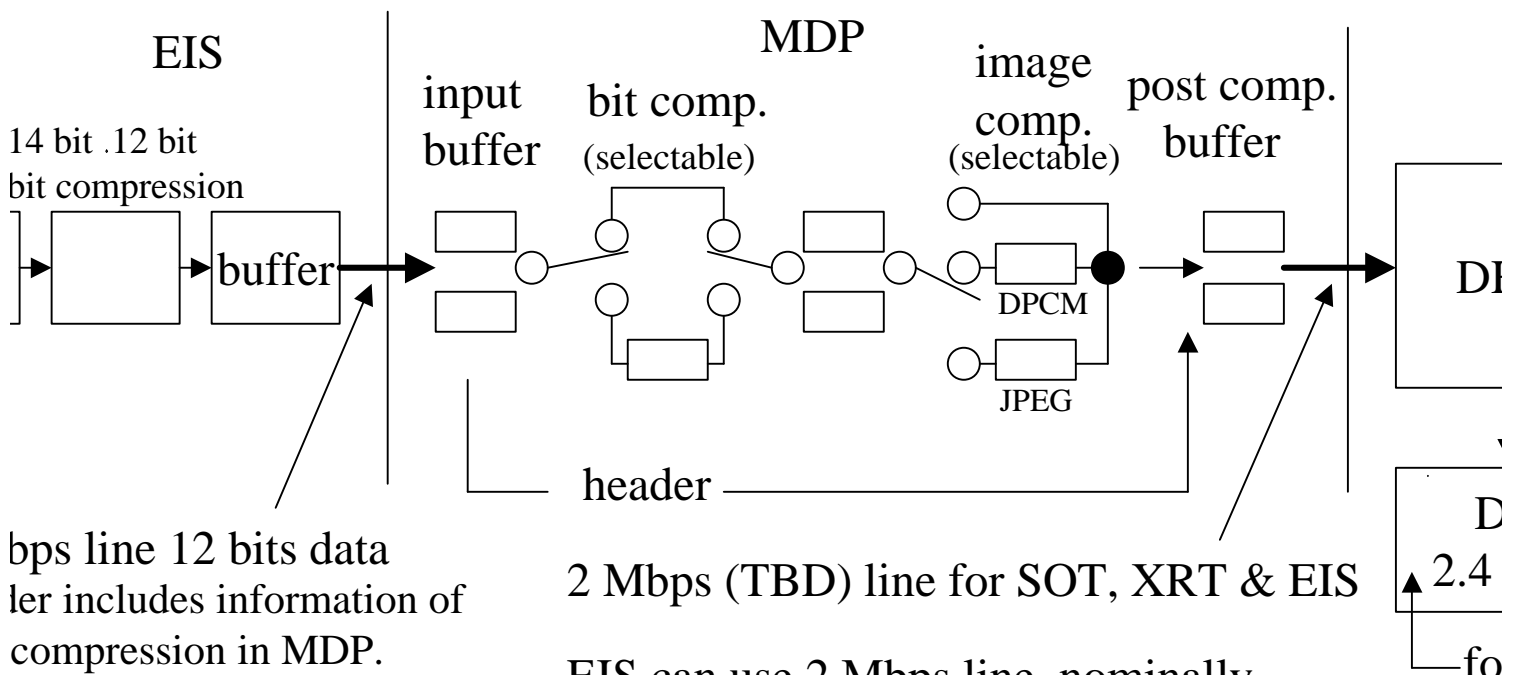
← Data format is consistent with requirements

REQUIREMENTS

Request of sending status data is informed from MDP via command line, and EIS will send status data to MDP within 0.5 sec.







bps line 12 bits data  
ler includes information of  
compression in MDP.

EIS can use 2 Mbps line nominally  
for 0.5 s every 6 seconds.

→ max. 167 kbps, but DR for EIS will  
become full in a short time.

y of DR will be ~3 Gbits, but only downlink 2.4 Gbits during the KSC/DSN c

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## Command Lines

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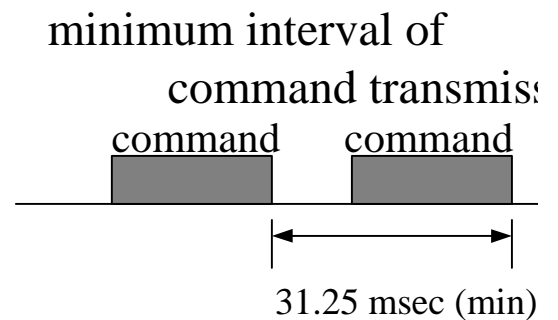
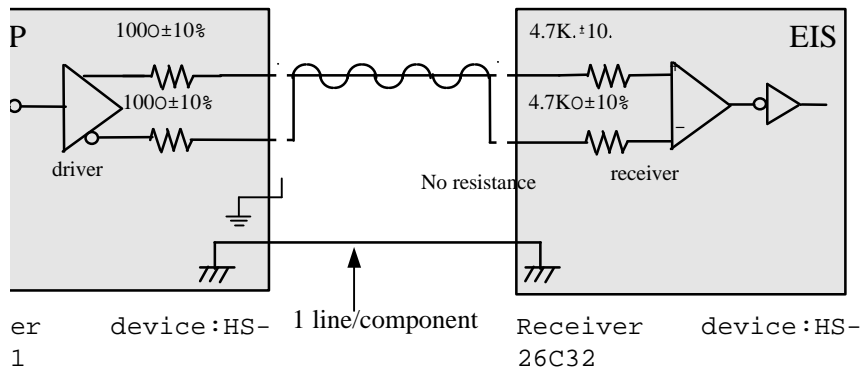
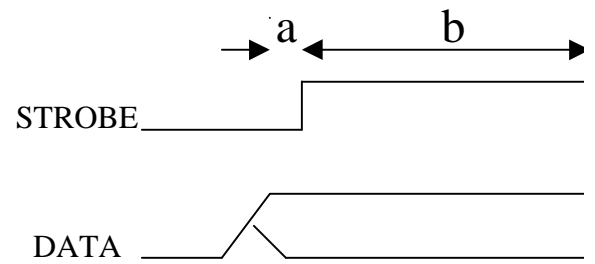
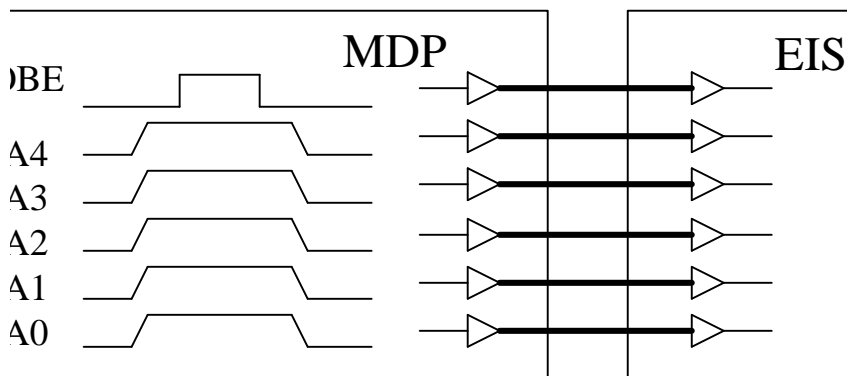
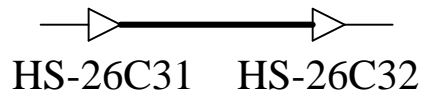
Command line: low-power differential RS-422 line, 64 kbps

Points contractors would like to know:

1. list of commands required for EIS operation
2. how ICU reads commands from MDP (Please explain it.)
3. response time for a command
4. size of buffer ?
5. whether there is some restriction for the data format or not.

Command answer back: receipt of command → put the command data into status for confirmation.

Purpose: Irrespective of MDP CPU status, the secondary power on/off command and others are sent via hardware command interface.

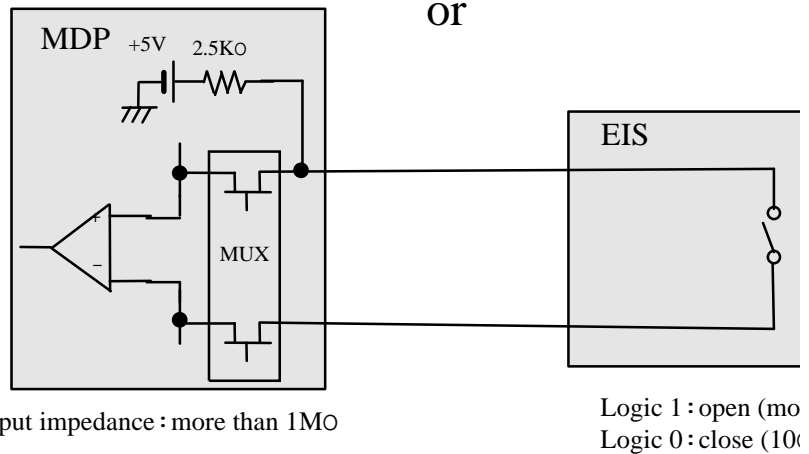
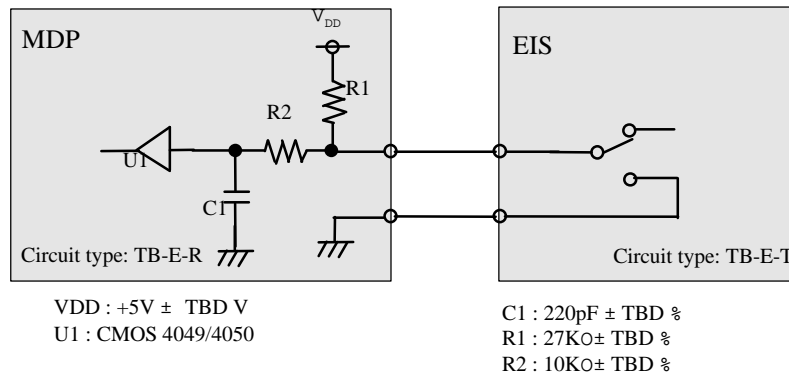
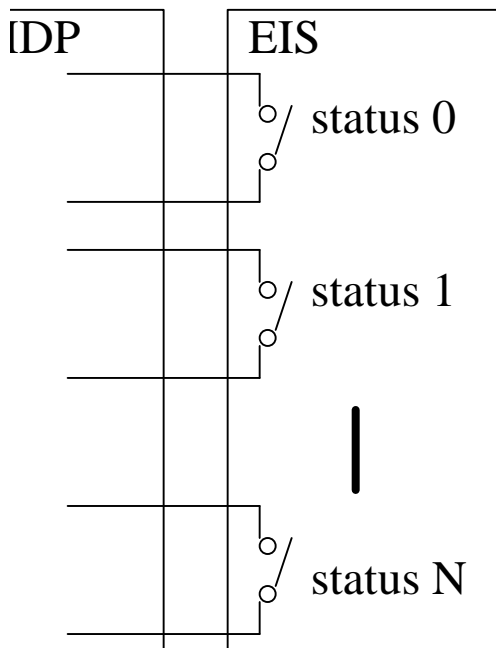


Source: Solar-B Electrical Design Standards § 8-(8)

# Solar-B

## Hardware status interface for secondary power

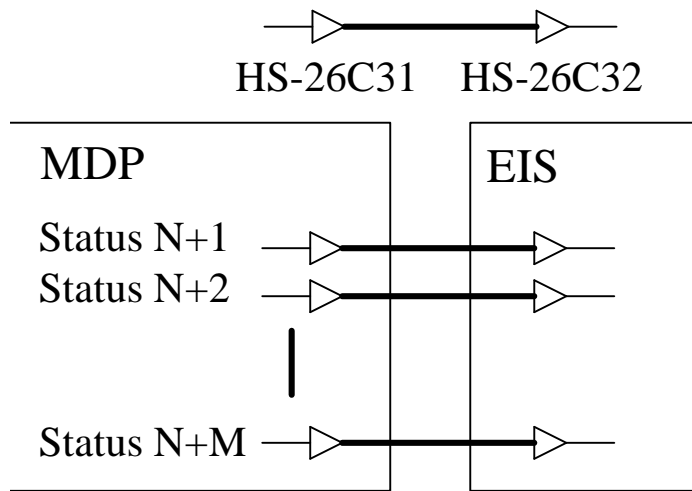
pose: ON/OFF status of EIS secondary power relay is periodically monitored a hardware logic circuits.



reference: Solar-B Electrical Design Standards §8-(6) or (7)

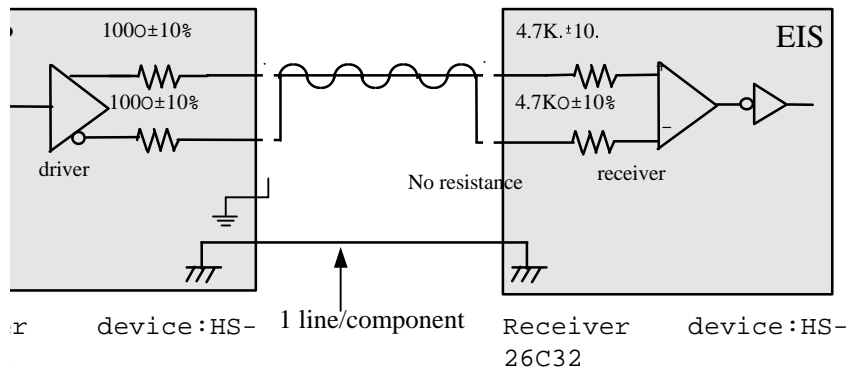
## Solar-B Hardware status interface for important bi-level s

Purpose: Important bi-level status data are periodically monitored by a hardware logic circuit.



N: Number of “secondary power

M: Number of “H/W status”



Reference: Solar-B Electrical Design Standards §8-(8)

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## Counter-plan for Errors

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When some error happens in EIS-CPU due to SEU,

does EIS side request anything to MDP ?

or

does EIS side deal with the error by itself ?

When some error happens in MDP due to SEU,

does EIS side request anything to MDP backup system for safety ?

or

does EIS side deal with the error by itself ?

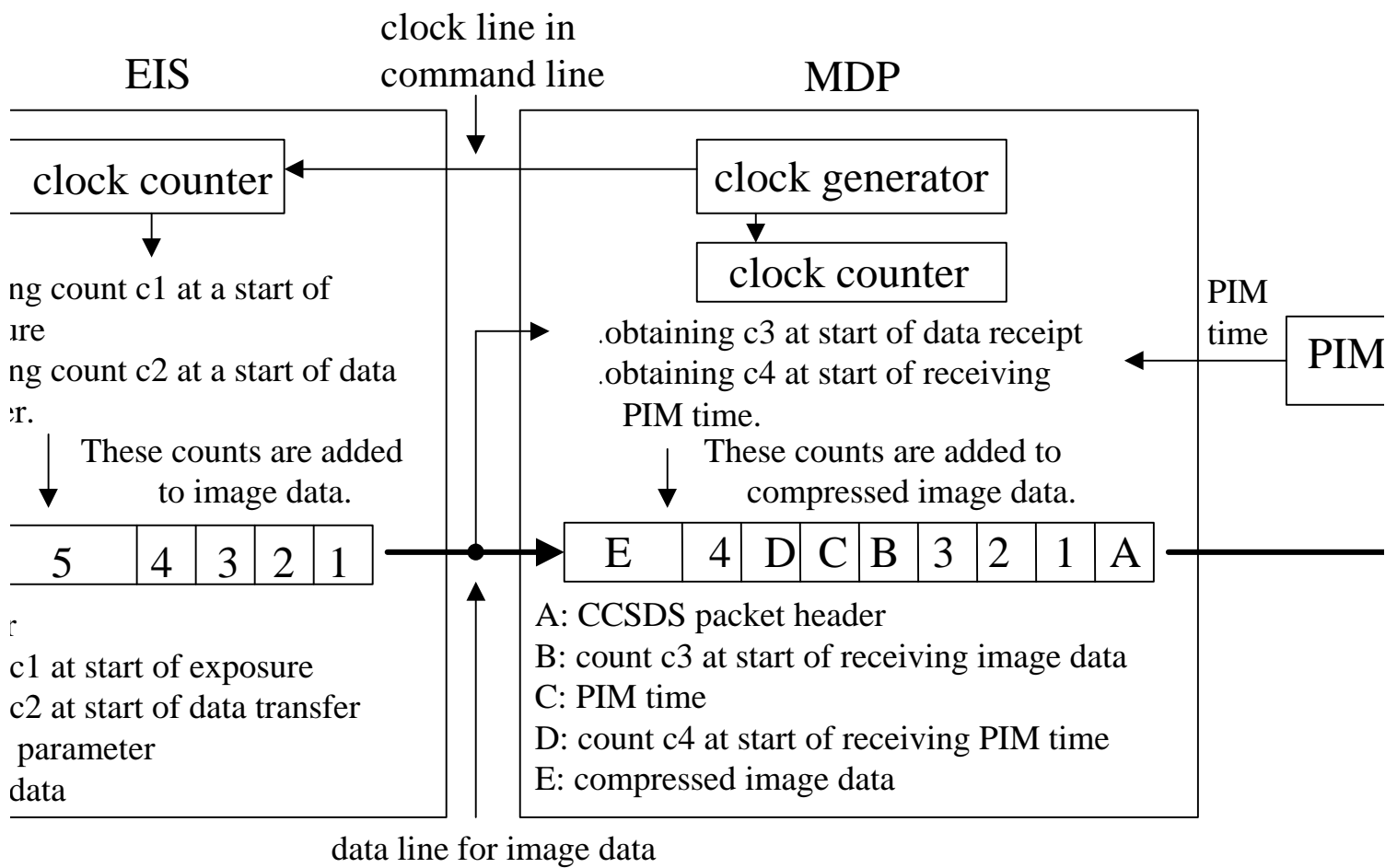
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MDP realizes stop of EIS CPU by the following procedure.

MDP sends a status request command and waits for 0.5 sec.

When the same procedure is done three times in a condition that status data do not come to MDP, MDP judges that EIS CPU is down.

this is effective only when the status line is independent of science data line.>



of data except for science data, EIS only put counter value at start of data transfer.

The following is a baseline of flare detection.

Flare detection is done by XRT.

A 8 on-chip summation image covering the whole XRT field of view will be used for flare patrol.

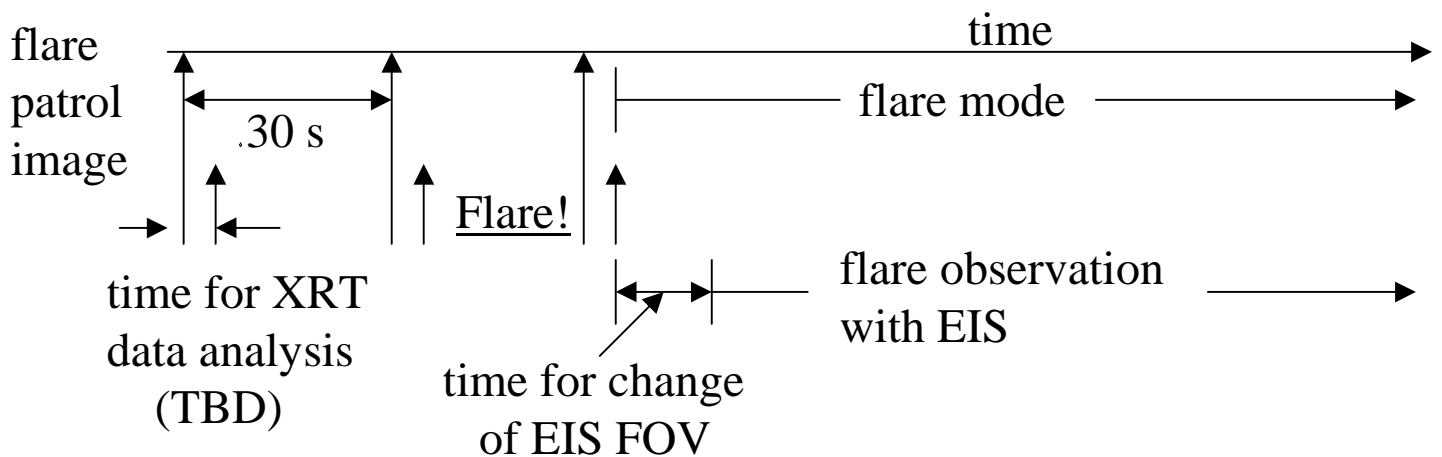
The flare patrol image will be taken every .30 sec (TBD).

Information on flare detection is sent to EIS by MDP.

1. Flare detection

2. Flare location in XRT CCD coordinate; X=0-255, Y=0-255 (TBD)

Duration of flare mode TBD



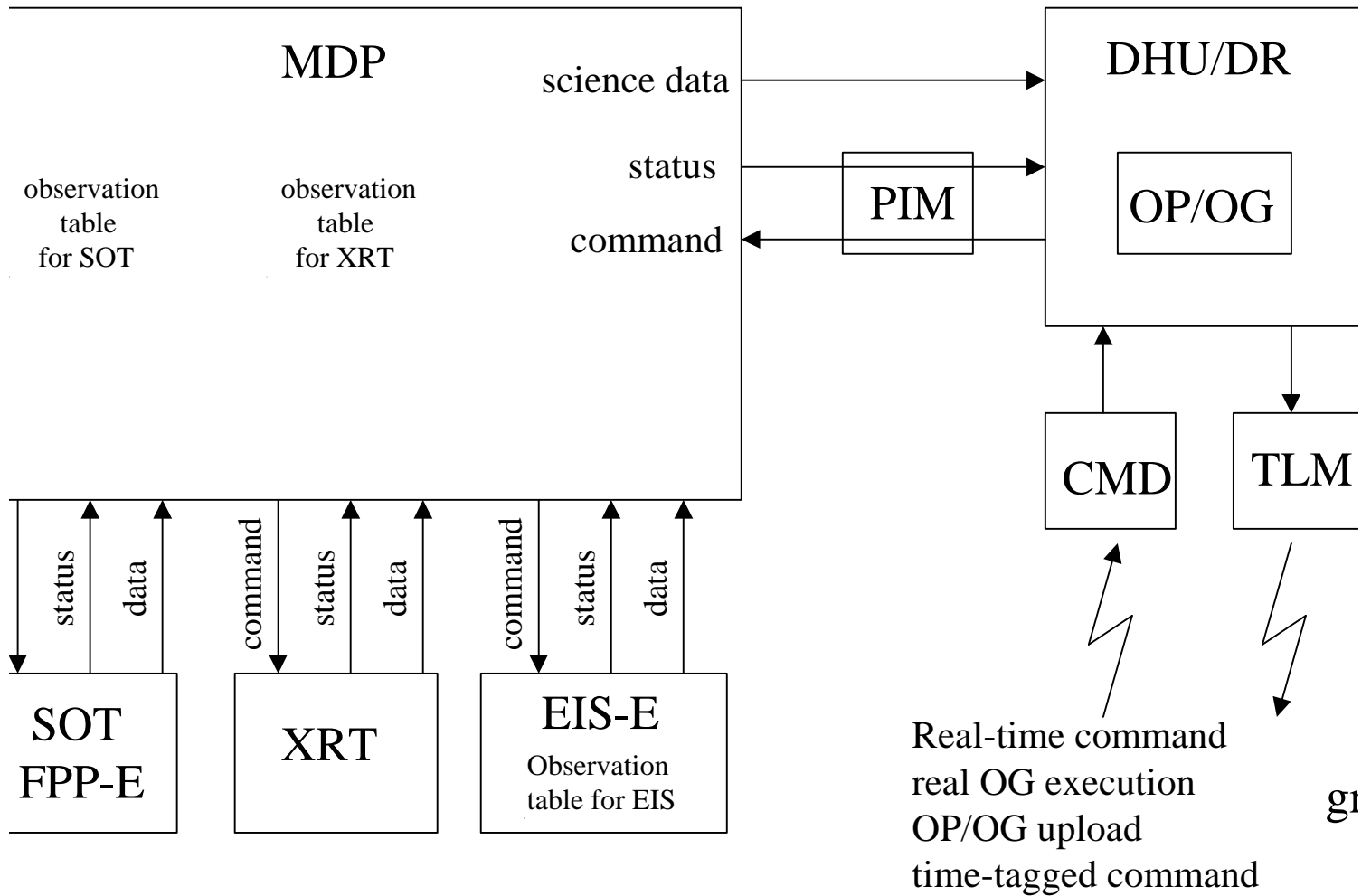


Solar-B	Command to MDP
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ASTRO-F type command system (similar to *Yohkoh* system) >

Type Classification	Command Types
Command Block Type	Discrete Command (DC) Block Command (DC+BC+...+ BC)
Command Grouping	Individual Command Organized Command (OG)
Operation Type	Real-time command Time-tagged command Operation Program (OP) Onboard-triggered command

OP and Onboard-triggered command support only OG to execute.



Solar-B		Data length of Command
Command Block Type		Maximum data length (bytes)
DC		1
DC+BC+.....+BC	real time	1 + 252
	OP/OG	1 + 13
	time-tagged	1 + 8

DC is used for single task such as ON/OFF, START/STOP and ENA/DIS.

BC is used for data transmission such as program load, table upload, and so on.

Real-time commands are sent from the ground and are immediately executed.

Time-tagged commands are stored in DHU and executed at a specified time in a resolution of about 1 sec. 32 commands can be stored in DHU.

**Operation Program;** a group of time-line commands consisting of OG. 512 (TBD) time slots are prepared for OP. Each time line is executed at a time defined by a time interval between consecutive two time lines. The unit of time is TBD sec. OP is started at a Kagoshima contact pass.

**Organized command;** One OG has 16 (TBD) command slots. Discrete command or block command data is included in each slot. 256 (TBD) OG can be stored in DHU. Time interval between two slots in an OG is 62.5 msec.

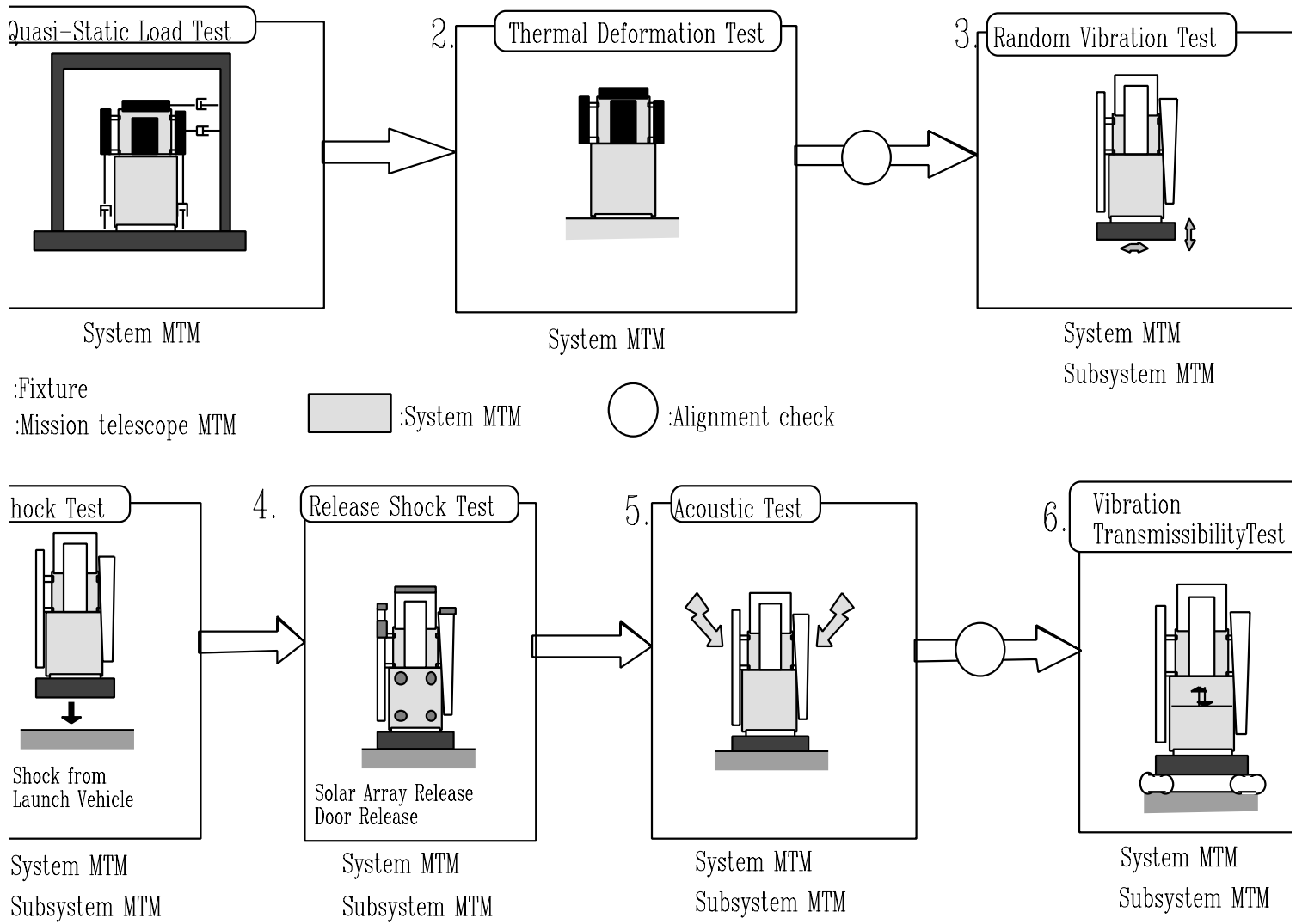
Example of OP:

CE	#OG	Interval	
00	0	121	NOP
01	3	32	KSC LOS
02	22	85	DSN AOS
.....			
129	255		OP STOP

Example of OG:

00:	03-DC-11	08:	05-DC-42
01:		09:	05-BC-12
02:	03-DC-23	10:	05-BC-45
03:		11:	05-BC-32
04:		12:	
05:		13:	
06:		14:	
07:		15:	03-DC-44

Test Flow



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## Mechanism & Operation

mechanism:	purpose	frequency	direction	angle of or sh
observation:				
	protect inside from dirty environment	once/year	ar. Y-axis	180deg/1
	increase of stiffness	once/mission	ar. X-axis	~30 deg
	focus adjustment	once/year	± Z-axis	±0.5 mm/10
	opening of vacuum enclosure	once/mission	ar. Y-axis	90 deg/1
	focus adjustment	once/mission?	± Z-axis	±0.5 mm/10
	change slit/slot	once/hour ?	ar. X-axis	90 deg/1
	change field of view	once/day ?	± X-axis	±5arcmin/10
observation:				
	scan for raster observation	once/ 1 sec	ar. Y-axis	1 arcsec/
	reset to start position	once/5-10 min	ar. Y-axis	8 arcmin
	adjust exposure duration	once/ 1 sec	ar. Z-axis	~60 deg/
	change slit/slot	once/ hour	ar. X-axis	90 deg/

Disturbance torque during a motion:  $T = d(I \mathbf{w})/dt \sim I \mathbf{q} (dt)^{-2}$

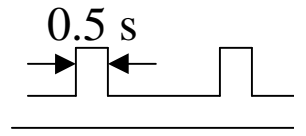
momentum of inertia  
estimated by CMB

unit (g cm<sup>2</sup>)

scan mirror:  $I_y = 4.89e4$   
 shutter :  $I_z = 36.7$   
 :  $I_x = 44.3$

scan mirror:  
[raster step]

$\mathbf{q} = 1$  arcsec,  $dt = 0.5$  sec



$T = 4.9e4 \cdot 10^{-7} \cdot 4.8e-6 \cdot 0.5^{-2} = 9.4e-8$  (N m)

[motion to home position]

$\mathbf{q} = 4$  arcmin,  $dt = 5$  sec

$T = 4.9e4 \cdot 10^{-7} \cdot 4.8e-6 \cdot 240 \cdot 5^{-2} = 2.3e-7$  (N m)

90 deg,  $dt = 10$  sec

$44.3 \cdot 10^{-7} \cdot 4.8e-6$   
 $\cdot 3600 \cdot 90 \cdot 10^{-2}$

$1.9e-4$  (N m)

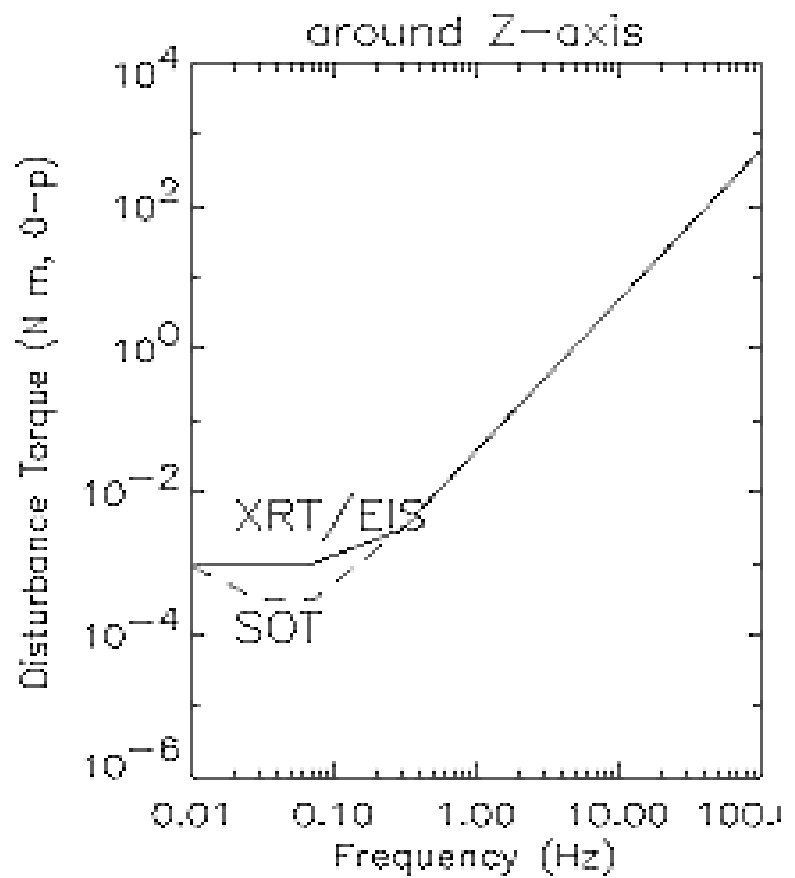
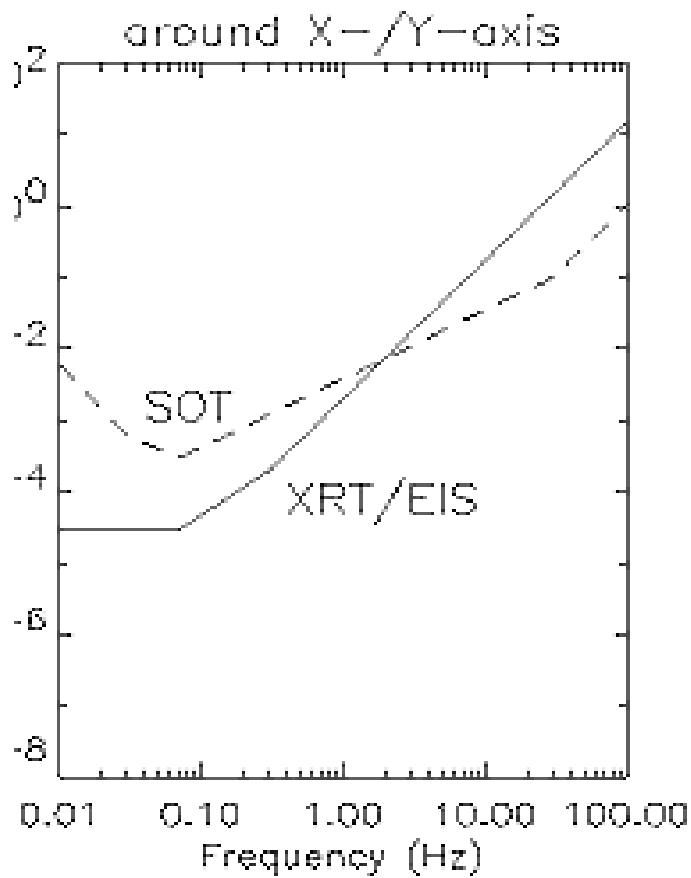
$6.9e-8$  (N m)

$6.9e-6$  (N m) for  $dt = 1$  sec

shutter:

$\mathbf{q} = 30$  deg,  $dt = 0.1$  sec

$T = 36.7 \cdot 10^{-7} \cdot 4.8e-6 \cdot 3600 \cdot 30 \cdot 0.1^{-2}$   
 $= 1.9e-4$  (N m)



from handouts distributed in satellite design meeting held on June 30, 1999